# KE<sub>T</sub> and Quark Number Scaling of v<sub>2</sub>

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#### Collaborated with

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#### Introduction

- Elliptic Flow (v<sub>2</sub>)
- Time Evolution

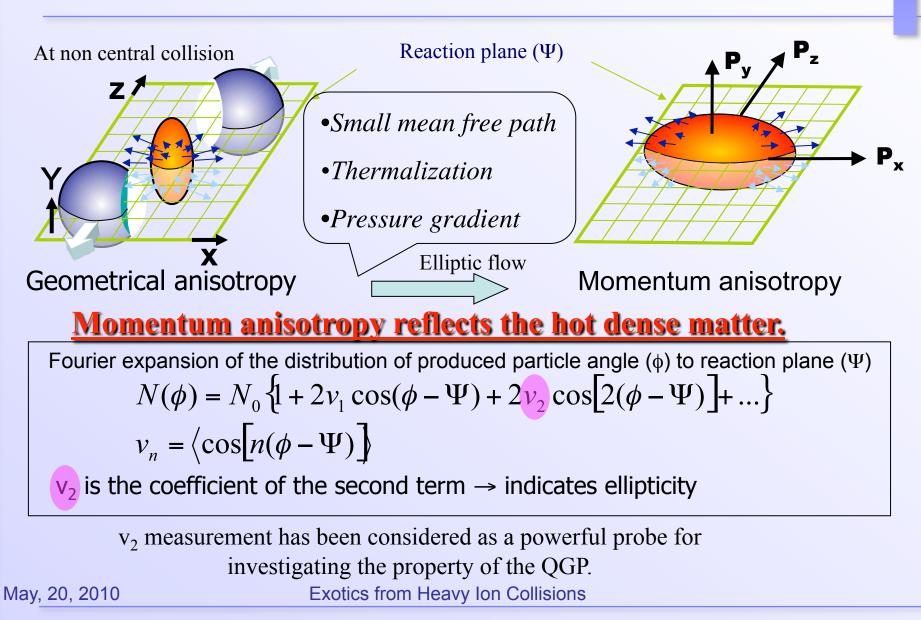
#### Results

- Fundamental Findings of v<sub>2</sub> at RHIC
  - Scaling of  $v_2$
- Blast-Wave Model Fit

Summary

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# Elliptic Flow (v<sub>2</sub>)



#### **Time Evolution** The matter produced in the high energy heavy ion collision is expected to undergo several stages from the initial hard scattering to the final hadron emission. **Kinematical freeze-out** Freeze-Out Hadron gas Chemical freeze-out Mixed phase Hadronization Hadron Gas Expansion & Cooling QGP Thermalization pre-equilibrium $\tau_0 \leq 1 \text{ fm/}c$ Hard scatterings Z Collision

When the matter is thermalized, we expect *Hydro-dynamical behavior at quark level*.

Note whenever the matter interacts each other,  $v_2$  could change.

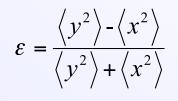
Need a comprehensive understanding from thermalization through hadronization to freeze-out. May, 20, 2010 Exotics from Heavy Ion Collisions

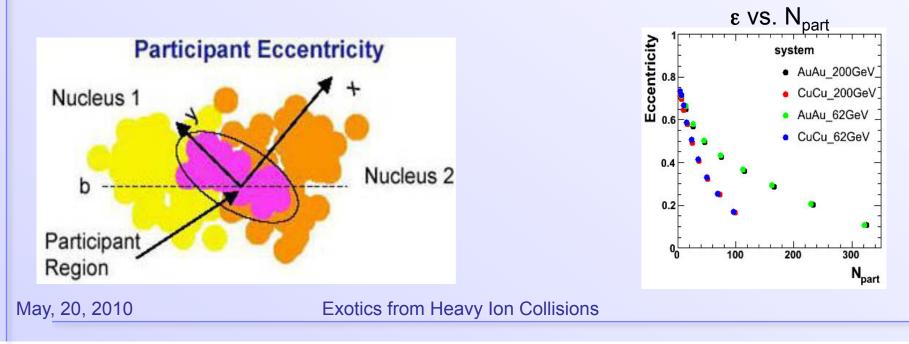
## Words

#### N<sub>part</sub> --- Number of nucleons participating the collision Eccentricity (ε) --- geometrical eccentricity of participant nucleons

- Monte-Carlo simulation with Glauber model
- Nucleus formed by wood-Saxon shape

- Participant eccentricity which is calculated with long and short axis determined by distribution of participants at each collision.



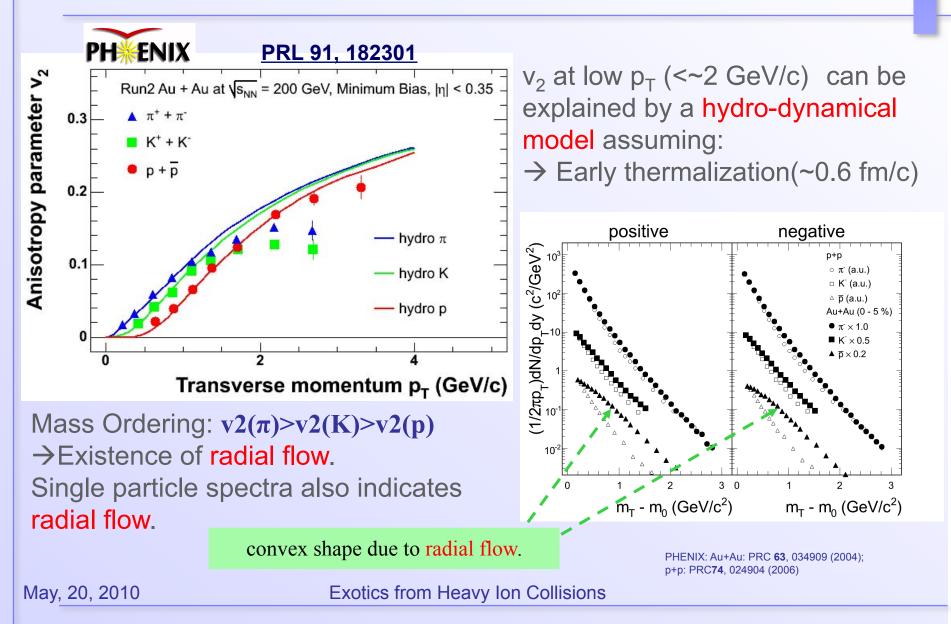


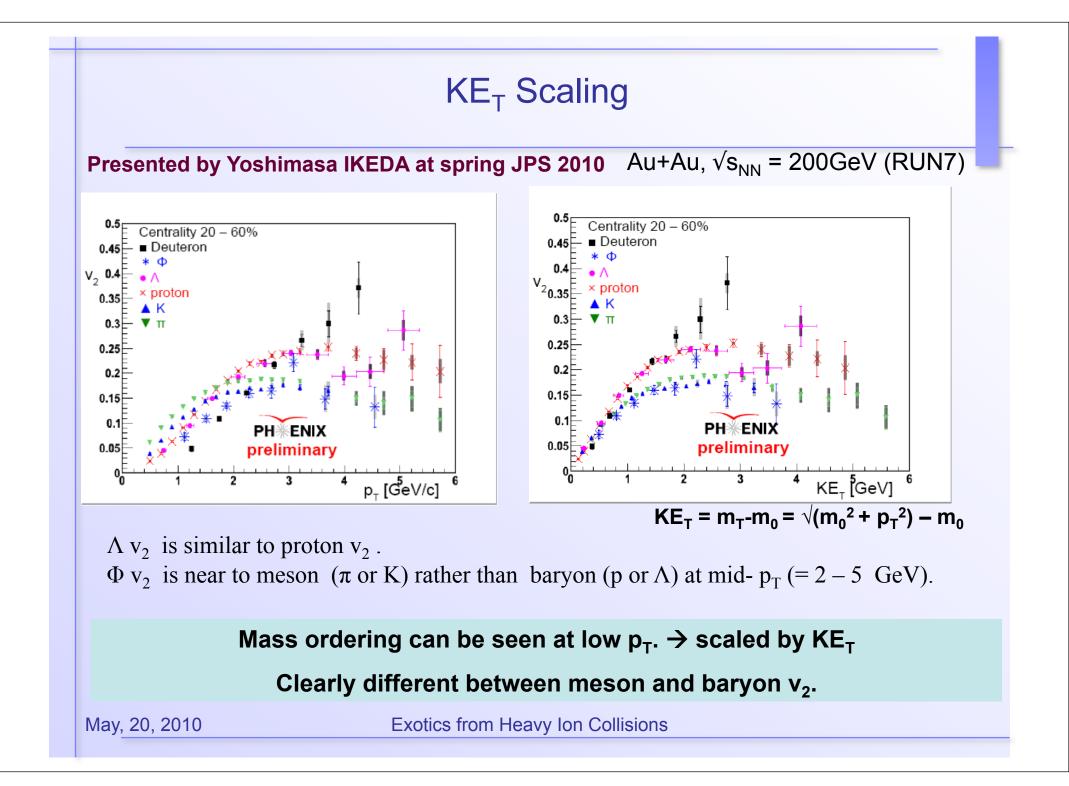
# Fundamental Findings of v<sub>2</sub> at RHIC

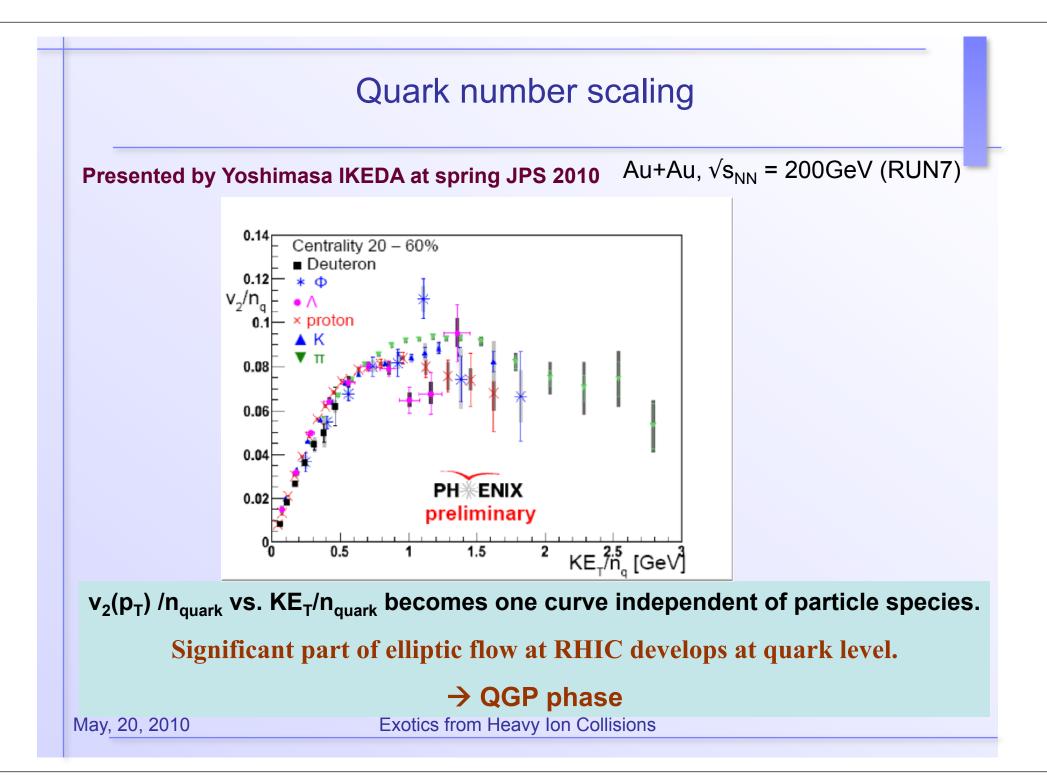
Hydro-dynamical behavior
KE<sub>T</sub> scaling
Quark number scaling

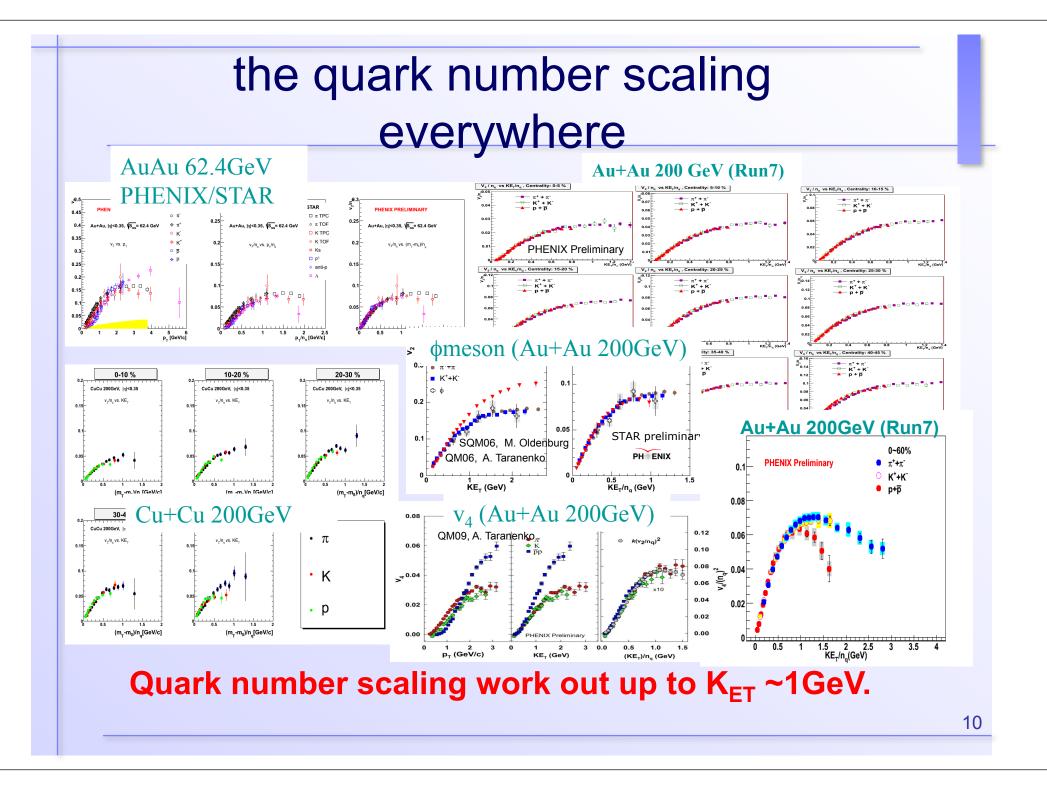
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# v<sub>2</sub> explained by hydro model



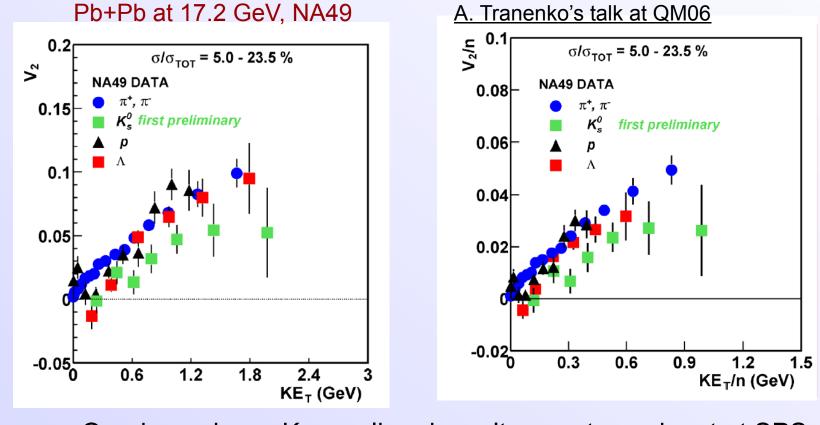






# quark number scaling at SPS

 $v_2$  of p,  $\pi$ ,  $\Lambda$  - C. Alt et al (NA49 collaboration) nucl-ex/0606026 submitted to PRL  $v_2$  of K<sup>0</sup> (preliminary) - G. Stefanek for NA49 collaboration (nucl-ex/0611003)



- Quark number + K<sub>ET</sub> scaling doesn't seem to work out at SPS.
- No flow at quark level due to nonexistence of QGP ?

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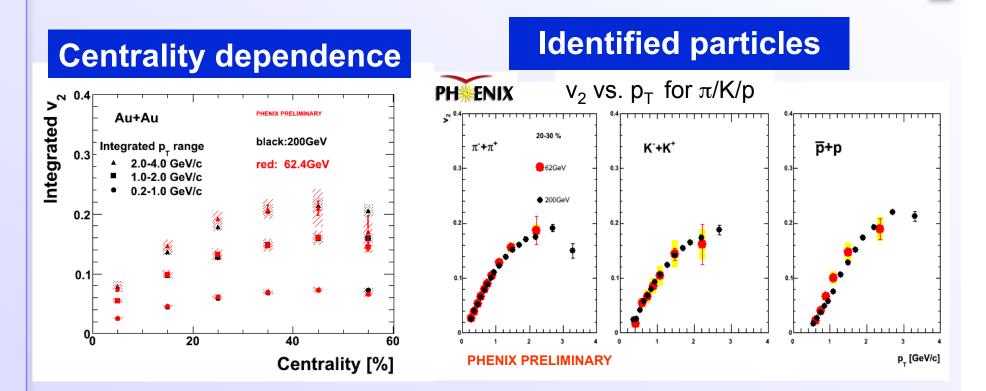
For a comprehensive understating of the matter and the mechanism of v2 production...

# Other scaling of v<sub>2</sub>

- Energy dependence
- Eccentricity scaling
- Npart scaling

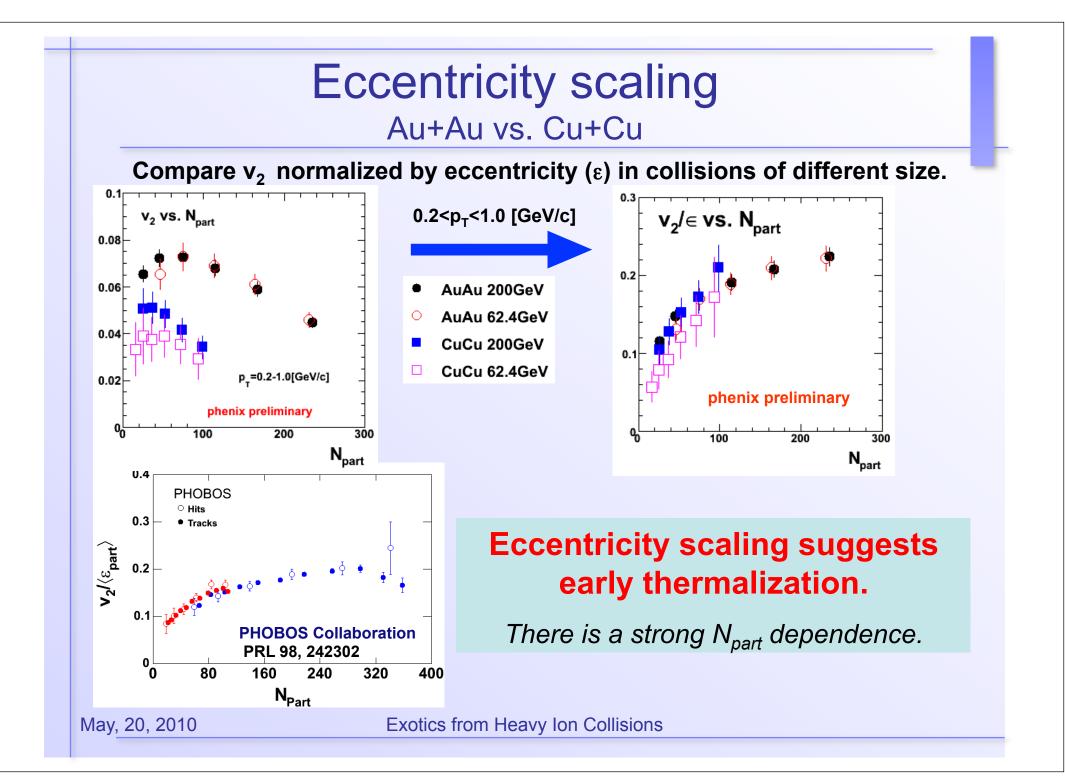
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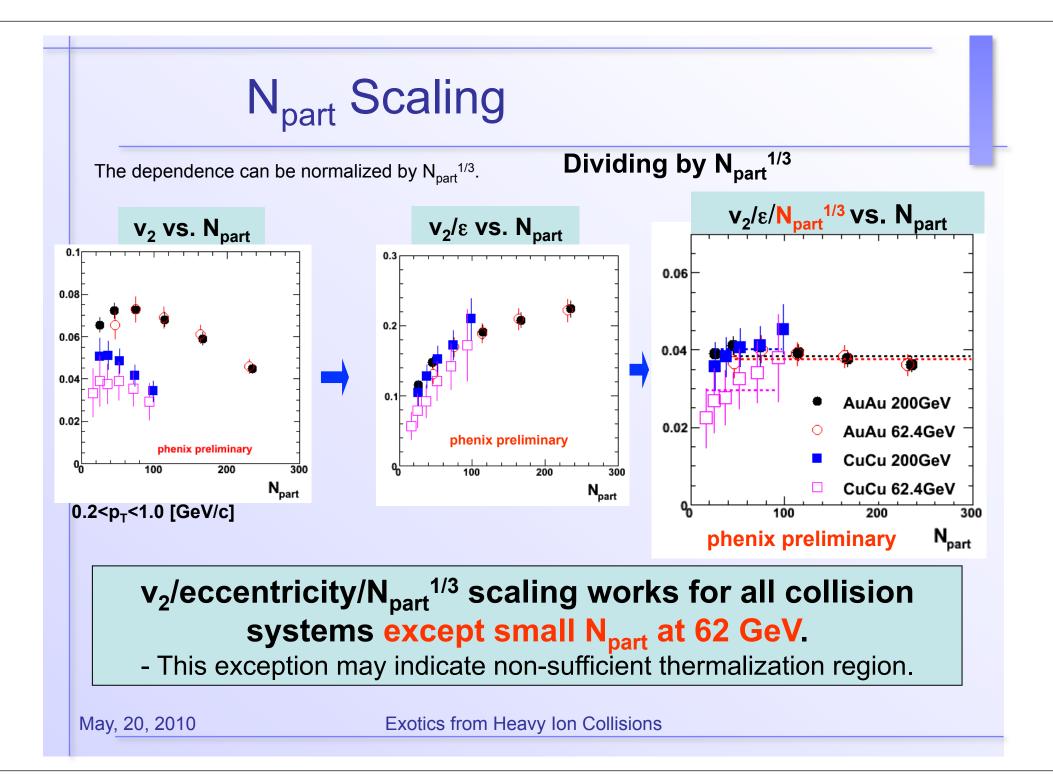
#### Energy dependence Au+Au 200 vs. 62 GeV



#### No significant difference between 200 and 62 GeV.

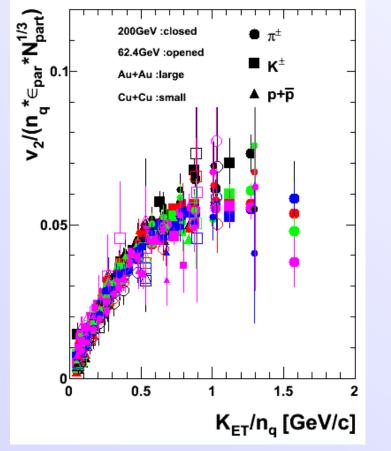
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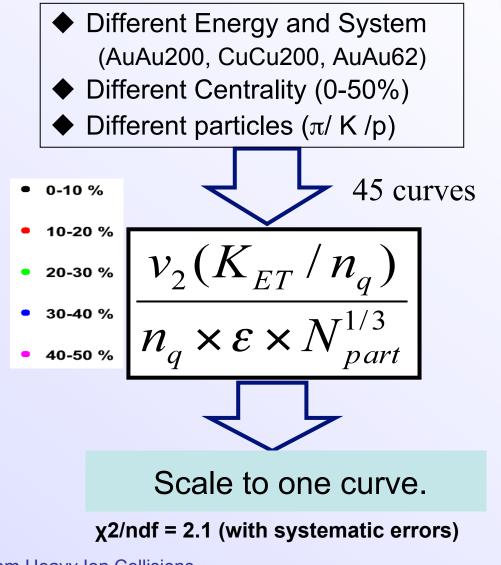




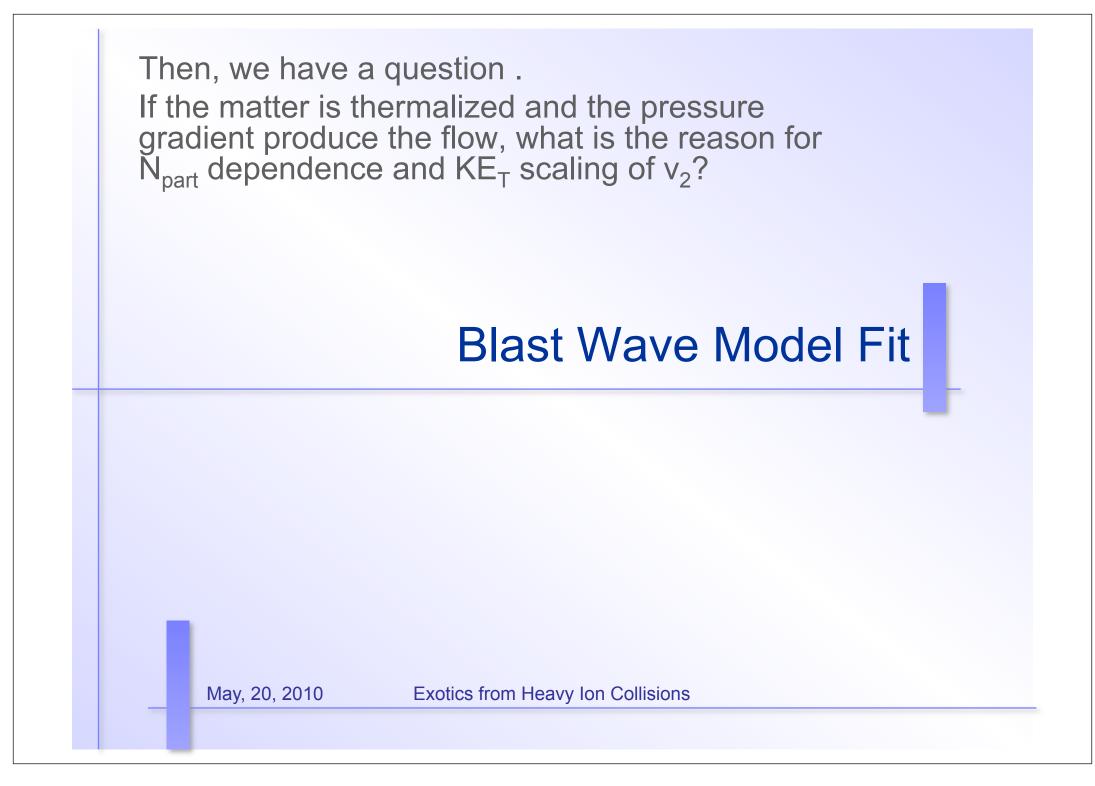
#### Universal v<sub>2</sub> for identified charged hadrons

Taking all scaling together,



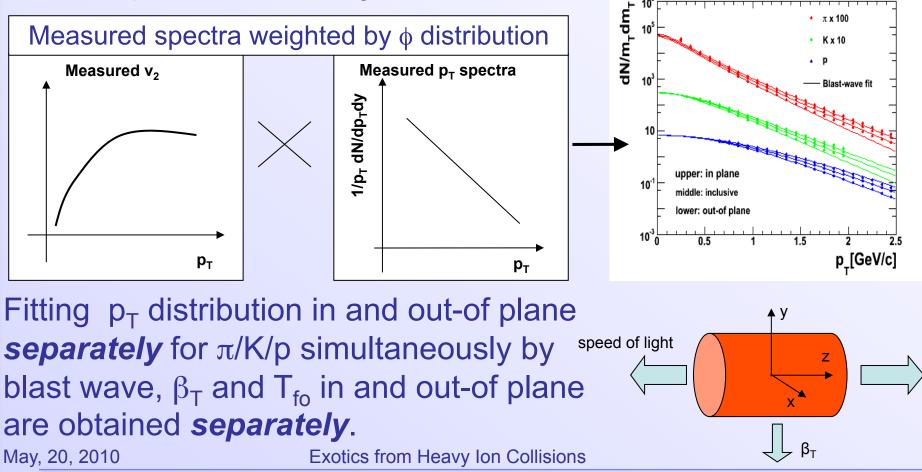


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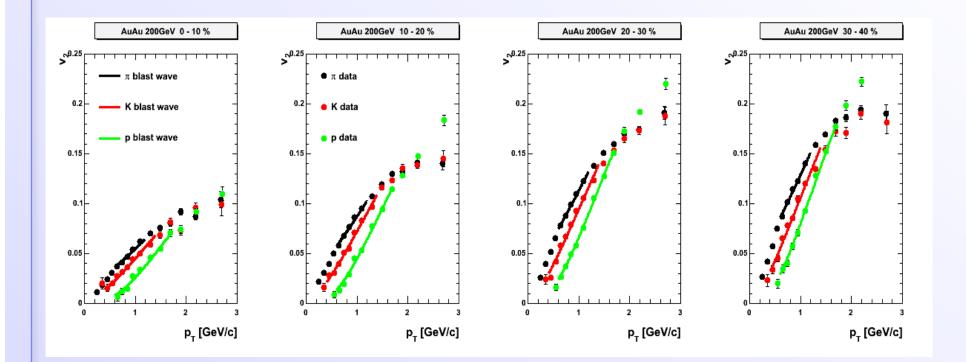


# Blast Wave Fitting for v<sub>2</sub> and Spectra

We use this well-known fitting technique to obtain the information of the flow velocity and temperature in and out-of plane **separately**.

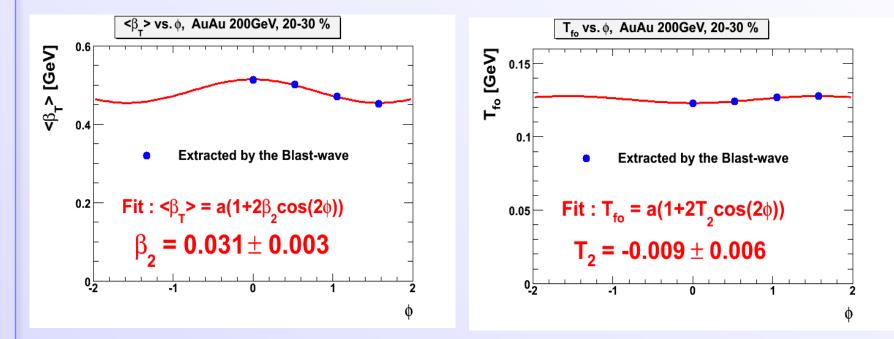


# Radial flow and $KE_T$ scaling



 Species dependence of v<sub>2</sub> can be reproduced by the Blast-wave model → Radial flow effect

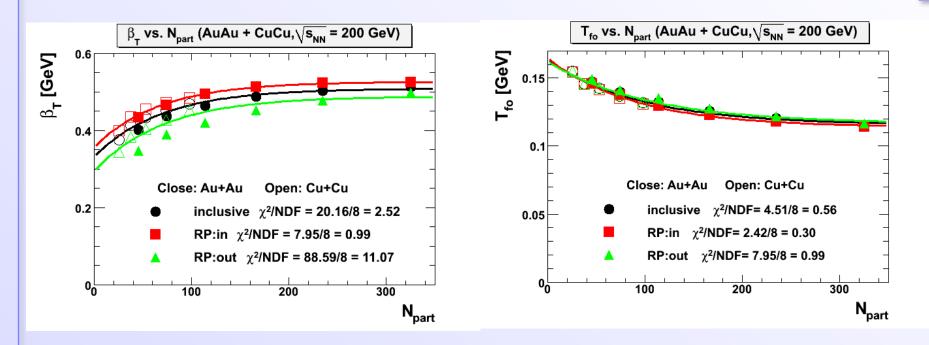
# Azimuthal dependence of $\beta_T$ and $T_{fo}$



- $\beta_T$  has clear azimuthal dependence.
  - Larger velocity @ in-plane
- T<sub>fo</sub> has small azimuthal dependence.
  - Lower temperature @ in-plane

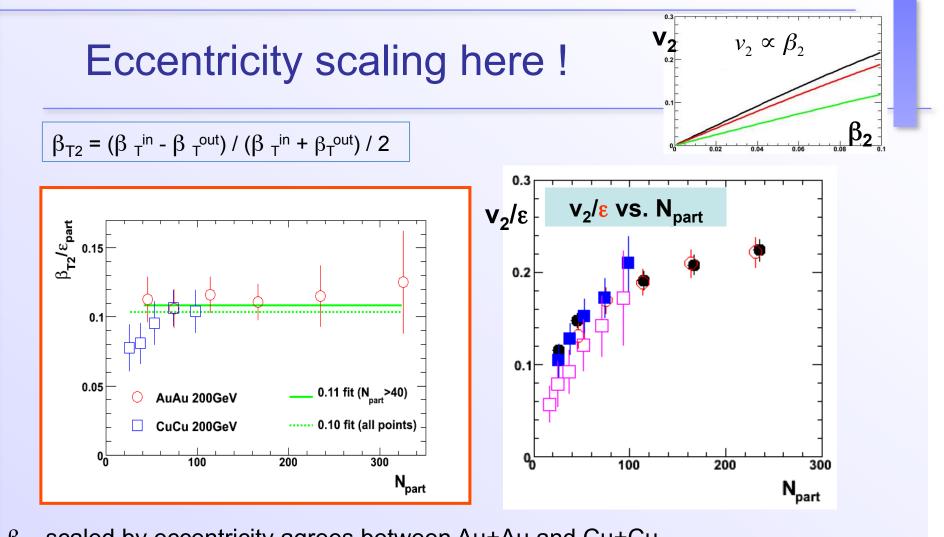
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# $N_{part}$ Dependence of $\beta_T$ and $T_{fo}$



 $T_{fo}$  and  $\beta_T$  agree between Au+Au and Cu+Cu, especially for the in-plane.

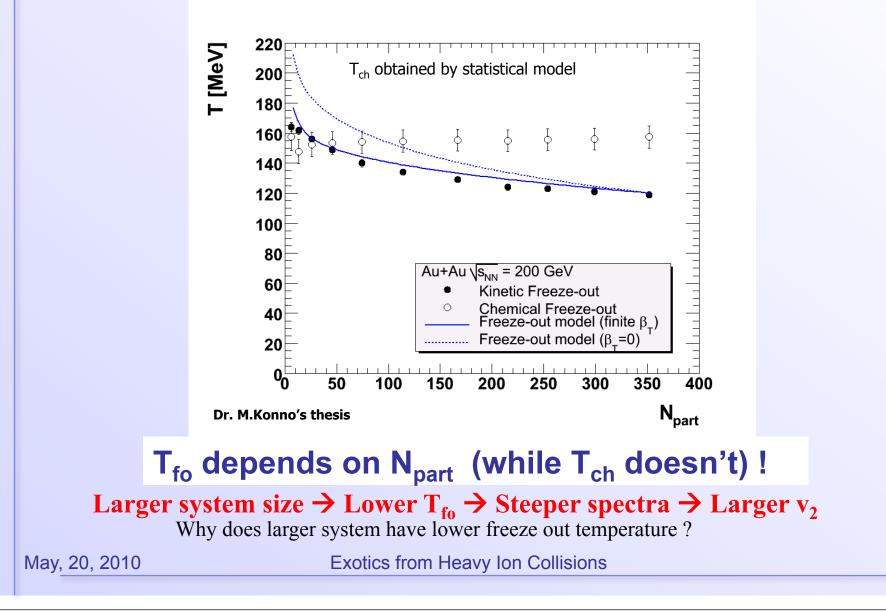
Since  $v_2$  is produced by the difference between in and out-of plane, the modulation of  $\beta_T$  is expected to have important rule to make  $v_2$ .



 $\beta_{T2}$  scaled by eccentricity agrees between Au+Au and Cu+Cu .  $\beta_{T2}$ /eccentricity is flat at  $N_{part} > 40$ .  $\rightarrow \epsilon$  drives  $\beta_{T2}$ !.  $\rightarrow$ Signal of Thermalization !?!?  $v_2$  is proportional to  $\beta_{T2}$  if other parameters are fixed. BUT,  $v_2$ / eccentricity is "not" flat  $\rightarrow$  What courses  $N_{part}$  dep. of  $v_2$  ??

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## Freeze-out Temperature and v<sub>2</sub>



## **Freeze-out Temperature and Time**

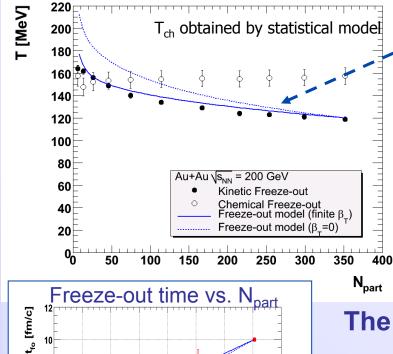
Dr. M.Konno's thesis

50

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100

150 200



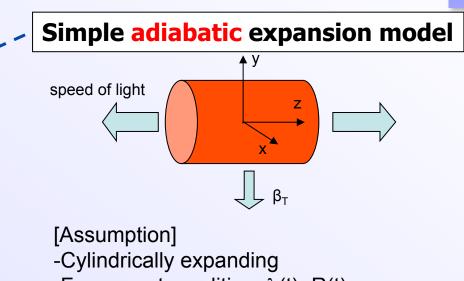
Au+Au \s<sub>NN</sub> = 200 GeV Adiabatic expansion

> Chemical Freeze-out Freeze-out model (finite  $\beta_T$ Freeze-out model ( $\beta = 0$ )

> > 350

400 N<sub>part</sub>

250 300



-Freeze-out condition:  $\lambda(t)=R(t)$ 

The model explains N<sub>part</sub> dependence well !

The times until freeze-out can be calculated by this model. Larger system takes more time to freeze-out.  $\rightarrow$ This makes lower T<sub>fo</sub>

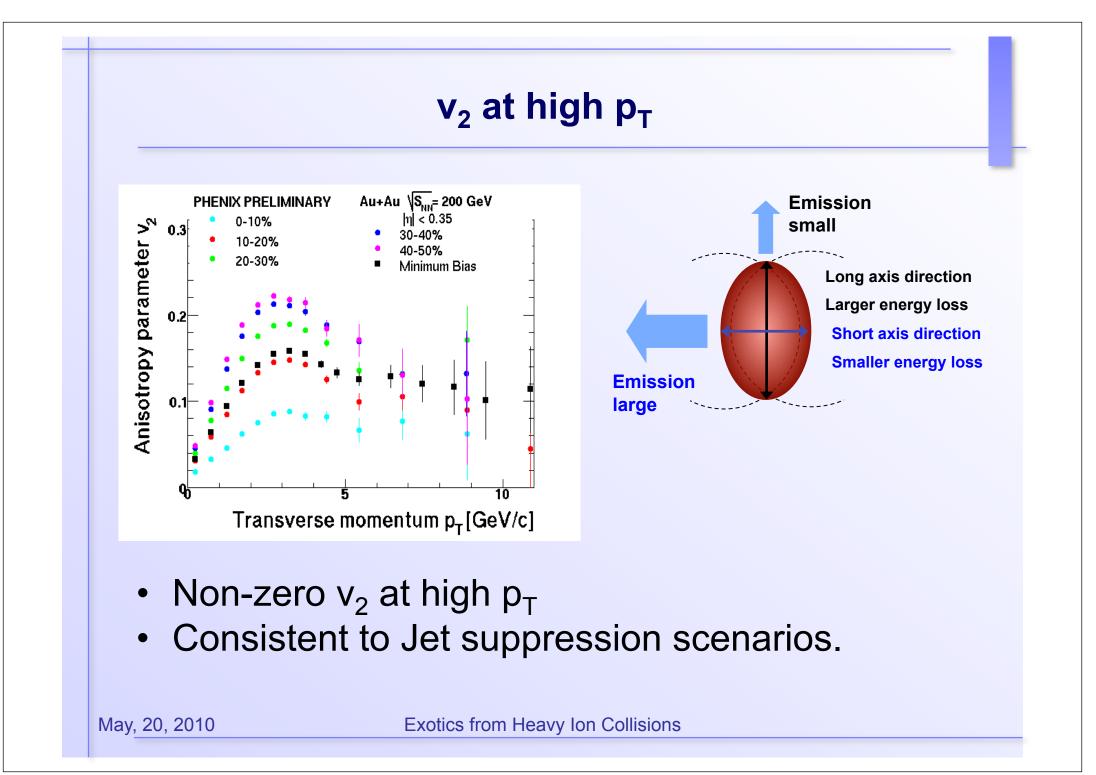
# Summary

- Systematic study of  $v_2$  have been done in Au+Au/Cu+Cu at  $\sqrt{s_{NN}} = 62.4/200$  GeV.
- v<sub>2</sub> values are saturated above 62.4 GeV in Au+Au.
  - Local thermalization
- $v_2(p_T)$  follows quark number +  $KE_T$  scaling in Au+Au (200,62GeV) and Cu+Cu (200GeV) .
  - Flow at quark level  $\rightarrow$  QGP phase
- $v_2(N_{part}) / \epsilon$  are same between Au+Au and Cu+Cu at 200 GeV.
  - Eccentricity scaling  $\rightarrow$  Early thermalization
- $v_2(p_T) / \epsilon / N_{part}^{1/3}$  scaling works except for small  $N_{part}$  at 62 GeV.
  - Existence of a universal  $v_2$  scaling at RHIC
  - Exception may indicate non-sufficient thermalization region.
- <From Blast-wave fit results with v<sub>2</sub> and spectra together>
- β<sub>2</sub>/eccentricity is constant not depending on system size (N<sub>part</sub>>40).
   Early thermalization !
- Larger system freezes out later at lower temperature.
  - cause the  $N_{part}$  dependence of  $v_2^{\prime}/~\epsilon$  .

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#### Energy dependence up to RHIC

FOPI : Phys. Lett. B612, 713 (2005). E895 : Phys. Rev. Lett. 83, 1295 (1999) CERES : Nucl. Phys. A698, 253c (2002). NA49 : Phys. Rev. C68, 034903 (2003) PRL 94, 232302 STAR : Nucl. Phys. A715, 45c, (2003). PHENIX : Preliminary. PHXENIX PHOBOS : nucl-ex/0610037 (2006) PHENIX pT (GeV/c)  $^{2}$ 0.20 STAR 1.75 (open) 0 In-plane CERES 0.65 (closed) E 895 Ō. 0.15 Centrality = 13 - 26 (%) 0.05 Δ 27 <sup>순 삼</sup> 0 ∑ <sup>0.10</sup> Ω Ο FOPI, charged E895, protons 0.05 E877 charged -0.05 CERES, charged NA49, pions 0.00 STAR, charged PHENIX, charged PHOBOS, charged Out-of-plane -0. -0.05  $10^{2}$ 10 100 10 ∕ู้∫ร<sub>่∧N</sub> (GeV)  $\sqrt{S_{NN}}$  (GeV)

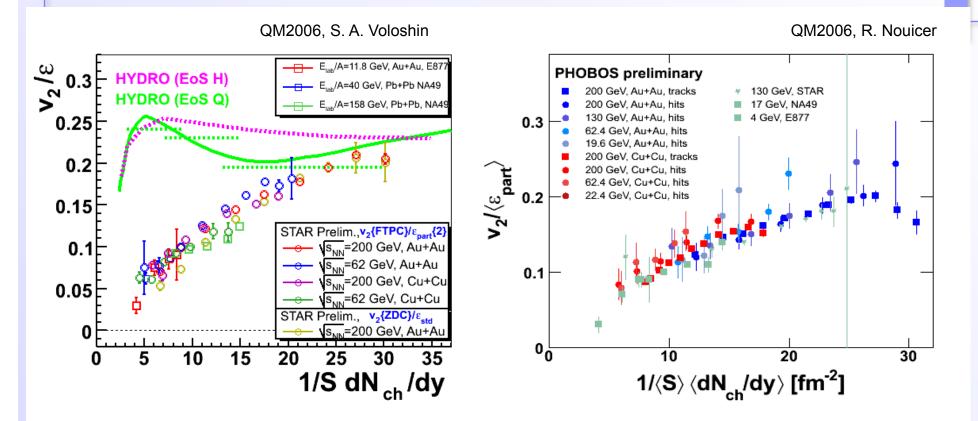
~ 50% increase from SPS to RHIC.

Above 62.4 GeV,  $v_2$  seems to be saturated.

→ The matter reaches thermal equilibrium state at RHIC.

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## Scaling (others)

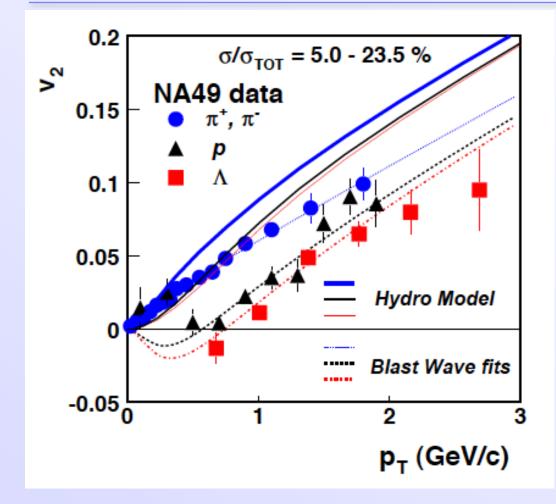


- Straight line from SPS to RHIC energy.
- v<sub>2</sub> is reaching the hydro limit at central collision ?

LHC and low energy scan may have answer for this !?

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#### v<sub>2</sub> compared with hydro model at SPS



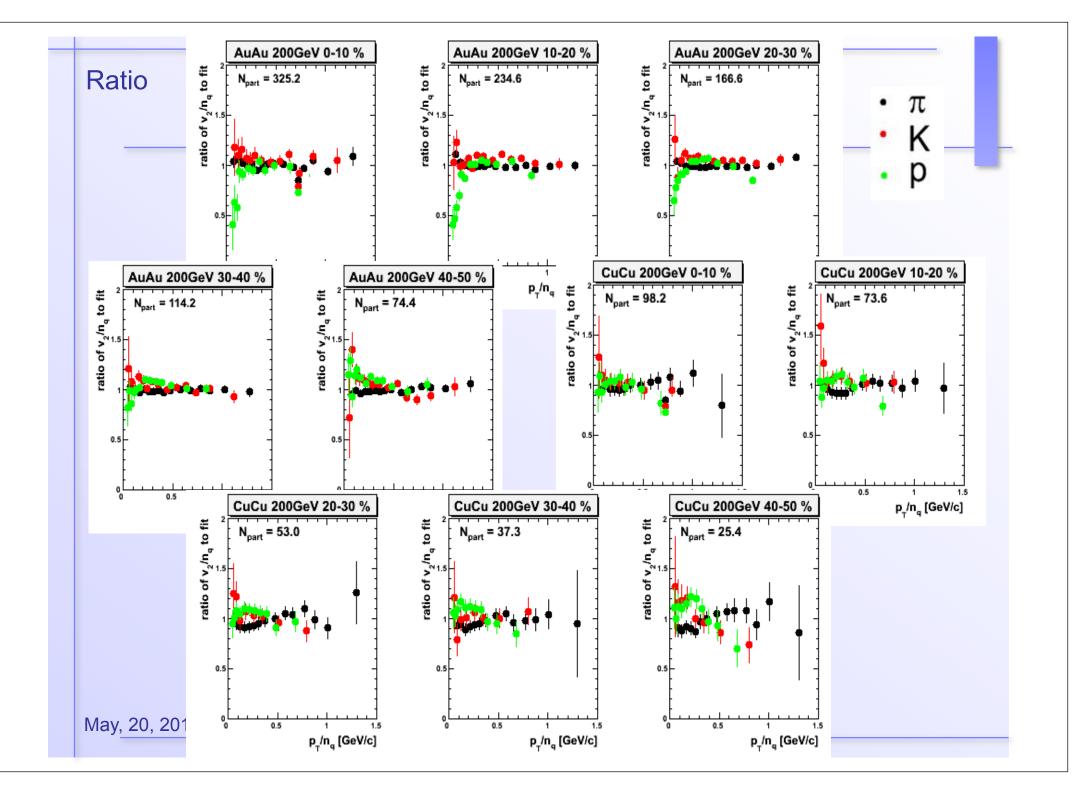
SPS ( $\sqrt{s_{NN}} = 17 \text{ GeV}$ )

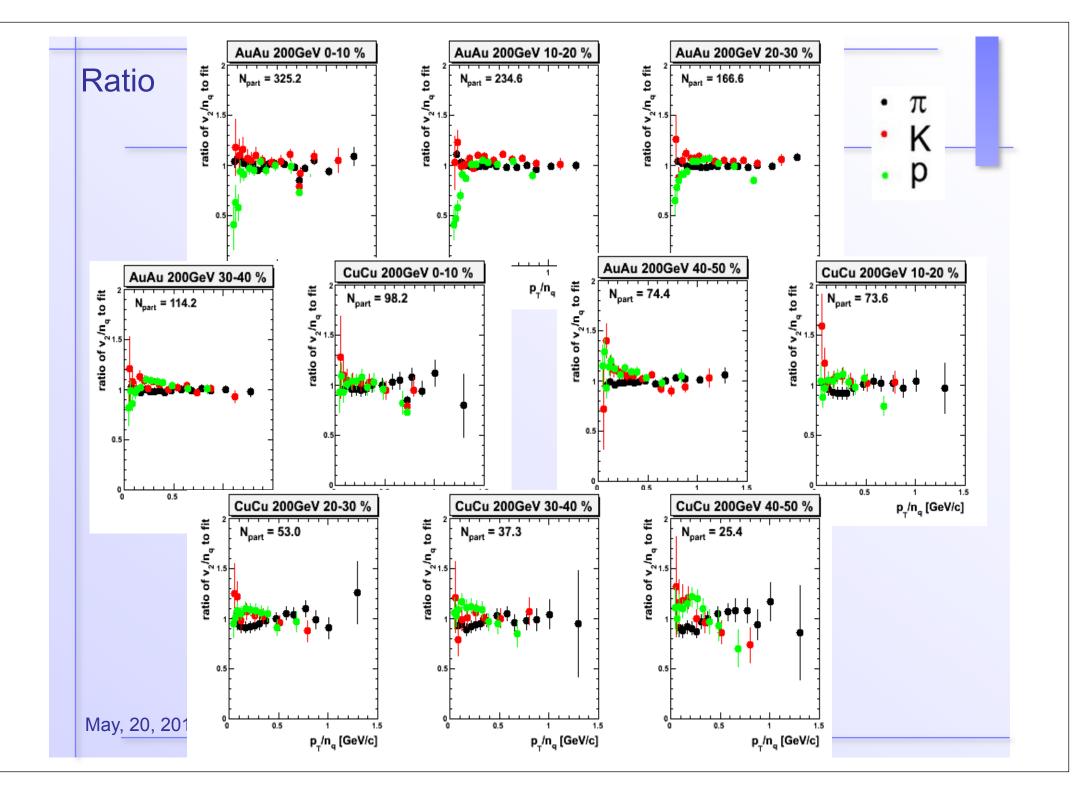
NA49: nucl-ex/0606026 (2007)

Hydro-dynamical model: 1<sup>st</sup> order phase transition, T<sub>c</sub>=165 MeV, T<sub>f</sub>=120 MeV,  $\tau_0$  = 0.8 fm/c

Hydro-dynamical model at SPS: Overestimate v<sub>2</sub>

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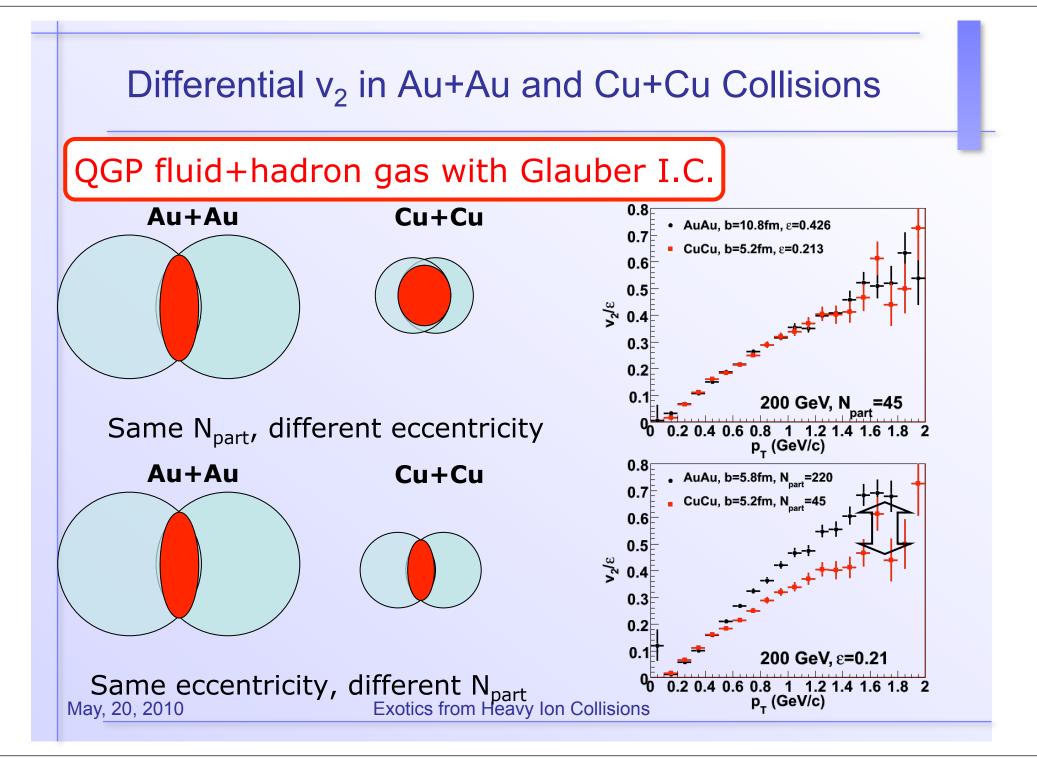


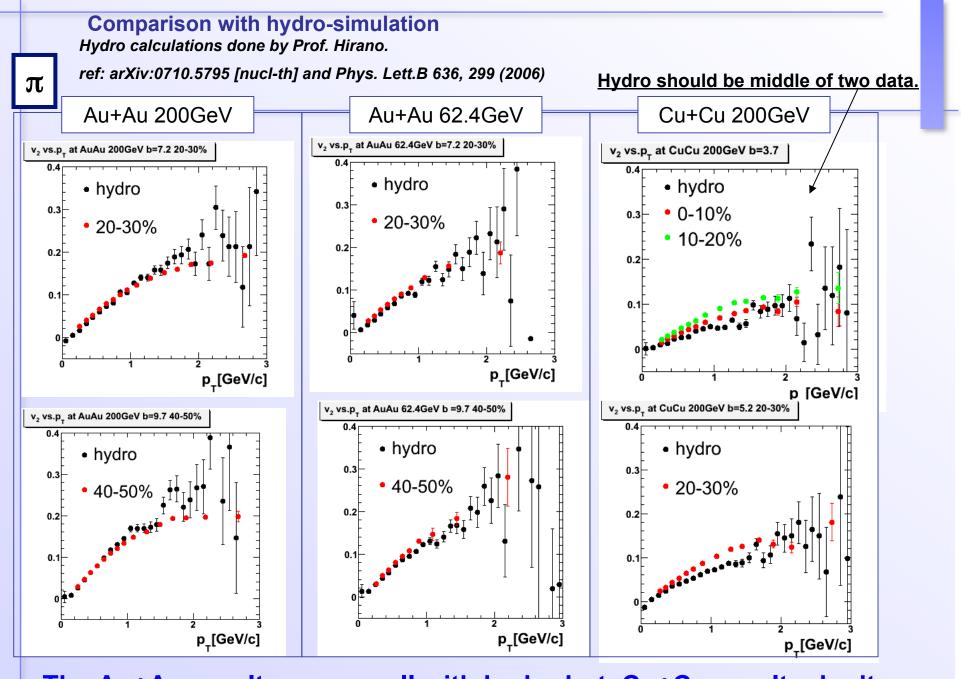


# Back Up

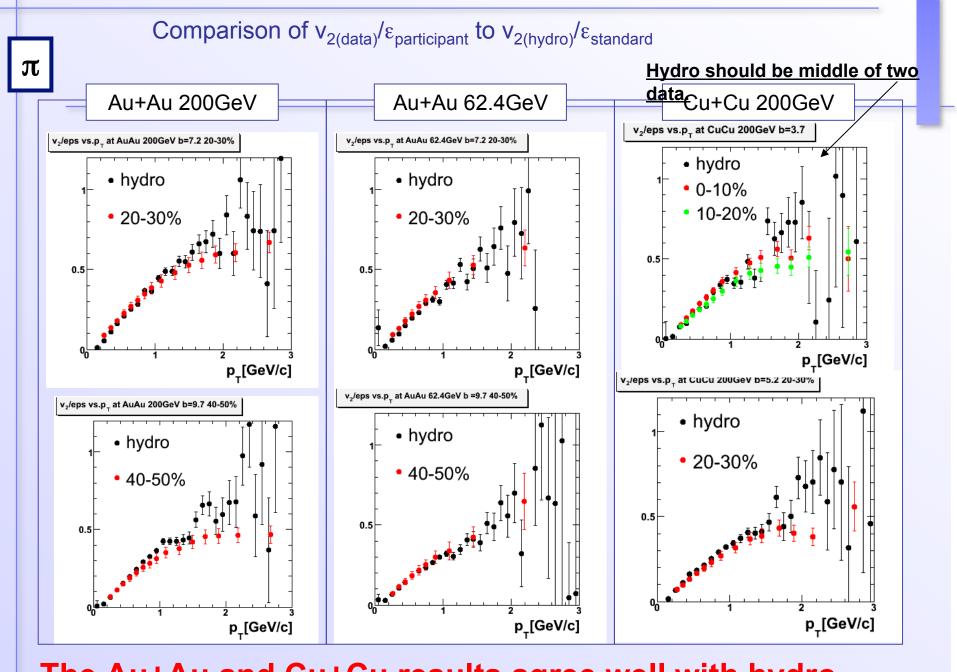
### **Comparison with Hydro simulation**

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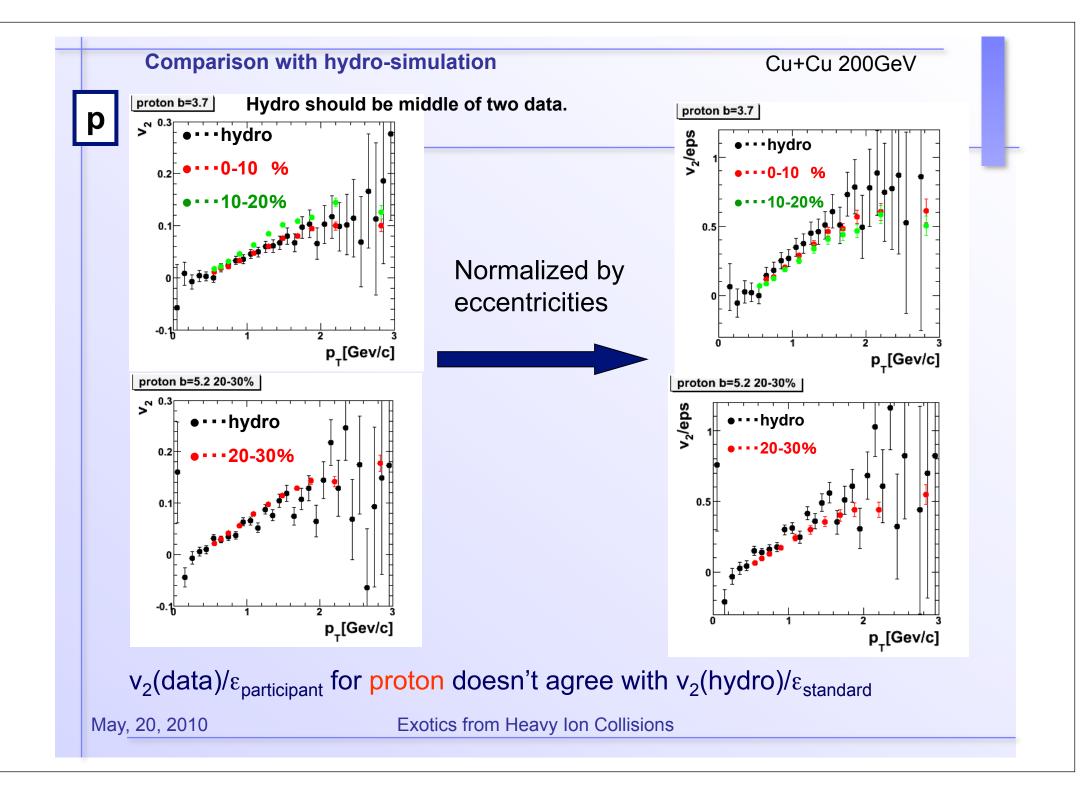


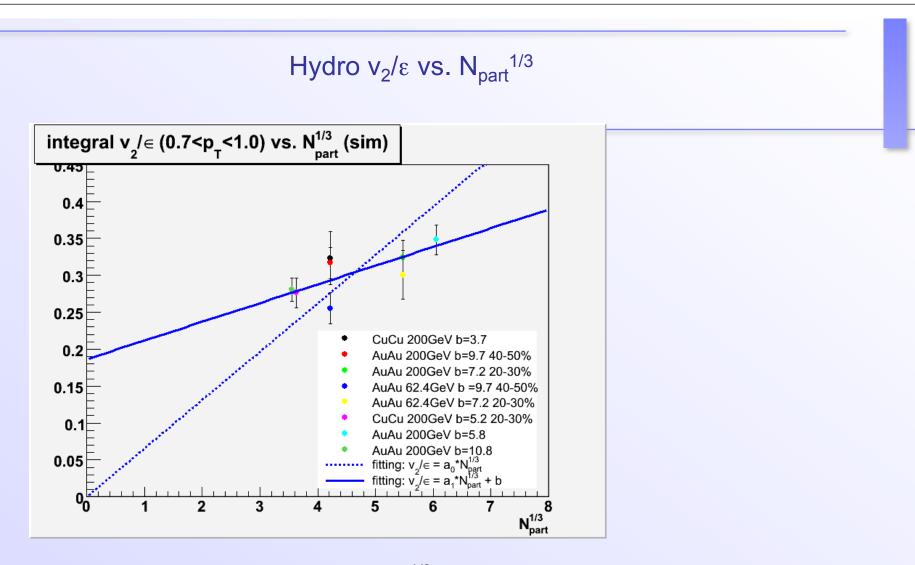


Mahe Auresults agree wellowith hydronisets Cu+Cu results don't.



MType Au+Au and Cu+Cu results agree well with hydro.



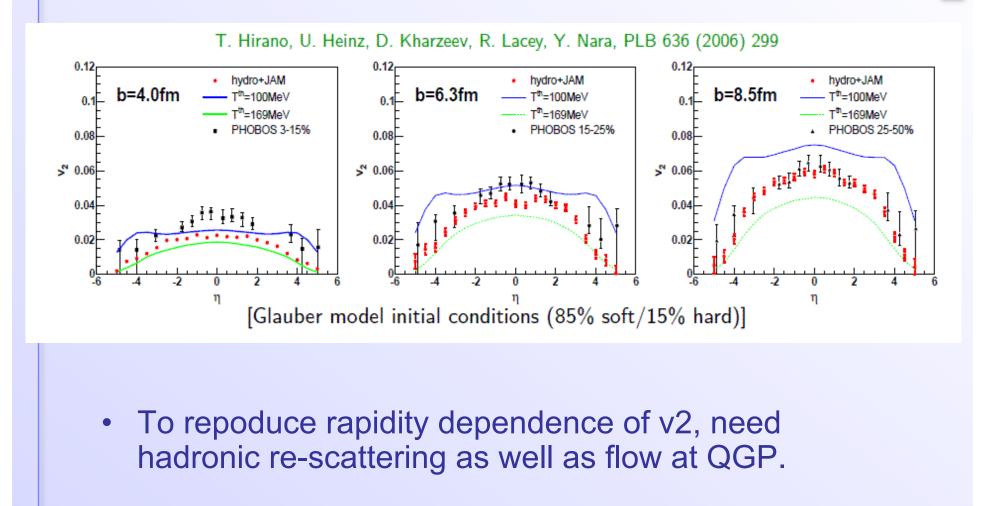


Fitting lines: dash line  $v_2/\epsilon = a^* N_{part}^{1/3}$ 

solid line 
$$v_2/\epsilon = a^* N_{part}^{1/3} + b$$

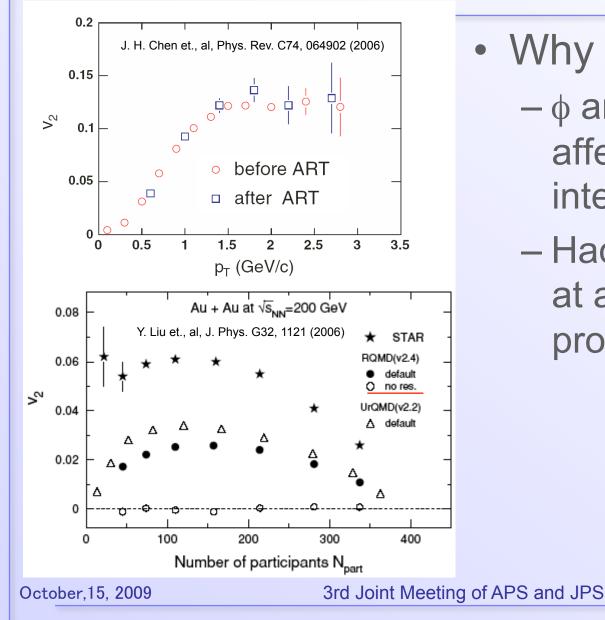
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#### Rapidity dependence



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## Multi-strange hadrons



- Why ?
  - $-\phi$  and  $\Omega$  are less affected by hadronic interactions
  - Hadronic interactions at a later stage do not produce enough  $v_2$